

# **Classical Optic Radial-Angular Entanglement**

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**Abstract** We study the classical optic entanglement between the radial and angular degrees of freedom in Laguerre-Gaussian mode superpositions with regard to symmetric first-order optical systems. The role of Gouy phase picked by a Laguerre-Gaussian mode on free propagation in regard to the radial-angular entanglement in the mode superpositions is brought out. We obtain a witness of radial-angular entanglement in two mode Laguerre-Gaussian superpositions which is demonstrated to be a robust free space signaler in the presence of atmospheric turbulence.

## Introduction

Paraxial light fields are solutions of the paraxial wave equation and retain their paraxiality on passage through symmetric first-order optical systems. Examples of such light fields are the Laguerre-Gaussian (LG) modes:

 $\Psi_{jm}(r,\theta) = \psi_{jm}(r;q) \times \Theta_m(\theta)$ , where  $\psi_{jm}(r;q) = \sqrt{\frac{2}{\pi w^2}} \left( \frac{(j-|m|)!}{(j+|m|)!} \right)^{\frac{1}{2}} \times \left( \frac{\sqrt{2}r}{w} \right)^{2|m|} L_{j-|m|}^{2|m|} \left( \frac{2r^2}{w^2} \right) \exp\left[ \frac{i\pi r^2}{\lambda q} \right],$ with  $\frac{1}{q} = \frac{1}{R} + \frac{i\lambda}{\pi w^2}$  and  $\Theta_m(\theta) = \exp[i2m\theta]$ 

# Witness of Radial-angular Entanglement

We define a measurable quantity, S which serves as a witness of radial-angular entanglement, which is also a possible free space signaler.

$$S = \left\langle \frac{r^2}{w^2(d)} \frac{-i\partial}{\partial \theta} \right\rangle - \left\langle \frac{-i\partial}{\partial \theta} \right\rangle \left\langle \frac{r^2}{w^2(d)} \right\rangle$$

For a two mode superposition,  $\Psi(r,\theta) = c_{j_1m_1}\Psi_{j_1m_1}(r,\theta) + c_{j_2m_2}\Psi_{j_2m_2}(r,\theta)$  $S_{j_2m_2}^{j_1m_1} = 2|c_{j_1m_1}|^2|c_{j_2m_2}|^2(j_1 - j_2)(m_1 - m_2).$ 

#### **Propagation through Atmospheric Turbulence**

Here  $L_{i-|m|}^{2|m|}(.)$  is the Laguerre polynomial with radial and azimuthal indices j and m, w is the width of the beam waist,  $\lambda$  is the wavelength, and q is the complex field parameter. For an LG mode propagating from the waist plane, we have

 $w^{2} \equiv w^{2}(d) = w^{2}(0) \left[ 1 + \left(\frac{d}{d_{r}}\right)^{2} \right], \text{ and } R \equiv R(d) = d \left[ 1 + \left(\frac{d_{r}}{d}\right)^{2} \right],$ 

where d is the distance of propagation from the waist plane, w(0) is the waist plane width, and  $d_r = \frac{\pi w(0)^2}{2}$ .

#### **Unitary Transformations**

Free propagation unitary transformation:

$$U_f(d) = \exp\left[-i\frac{d\lambda}{4\pi}\nabla_{\perp}^2\right], \text{ where } \nabla_{\perp}^2 = \frac{\partial^2}{\partial r^2} + \frac{1}{r}\frac{\partial}{\partial r} + \frac{1}{r^2}\frac{\partial^2}{\partial \theta^2}$$
  
Thin lens unitary transformation:

I nin lens unitary transformation:

$$U_l(f) = \exp\left[\frac{-i\pi}{\lambda f}r^2\right]$$

Action of free propagation unitary on an LG mode:  $U_f(d)\psi_{im}(r;a)\Theta_m(\theta) = e^{-i\phi_j(q;d)}\psi_{im}(r;a')\Theta_m(\theta)$ , where a' = a + d, and

$$\phi_j(q;d) = -(2j+1) \tan^{-1} \left(\frac{\frac{\lambda}{\pi w^2}}{\frac{1}{d} + \frac{1}{R}}\right)$$

Action of thin lens unitary on an LG mode:

$$U_l(f)\psi_{jm}(r;q)\Theta_{\rm m}(\theta) = \psi_{jm}(r;q')\Theta_{\rm m}(\theta),$$
 where

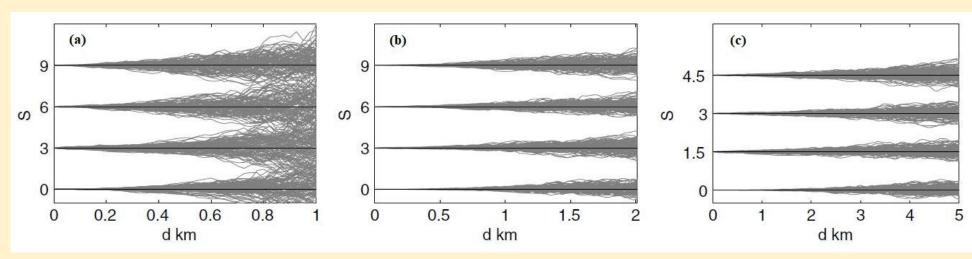
While both S and  $E(\Psi)$  remain invariant on free propagation for a two mode superposition, they need not be invariant as the superposition propagates through atmospheric turbulence. The phase spectrum of a two dimensional random phase can be written in terms of the Kolmogorov power spectral density as:

 $\Phi_{\theta}(K) = 2\pi \left(\frac{2\pi}{\lambda}\right)^2 \delta_d \Phi_n(K)$  where  $\delta_d$  is length of the interval, with  $\Phi_n(K) = 0.033 C_n^2 K^{-11/3}$ 

#### **Example:**

Consider the two mode LG superposition with mode indices  $j_1, m_1 = 0, 0$  and  $j_2, m_2 = \frac{3n}{2}, \frac{n}{2}, \text{ i.e.},$  $\Psi^{1n}(r,\theta) = \sqrt{(n-1)/n}\Psi_{00}(r,\theta) + \sqrt{1/n}\Psi_{\underline{3nn}}(r,\theta)$ (1) The witness of radial-angular entanglement for such a mode superposition is given by  $S_{\underline{3nn}}^{00} = \frac{3(n-1)}{2}$ .

# **Simulation Results**



**Fig. 2.** (a), (b) Witness of radial–angular entanglement S against distance of propagation d for

Any symmetric first-order optical system is realized as a sequence of free propagations and thin lenses.

# **Classical Optic Radial-angular Entanglement**

A paraxial field  $\Psi(r, \theta)$ , which is a finite superposition of LG modes, on Schmidt decomposition gives:

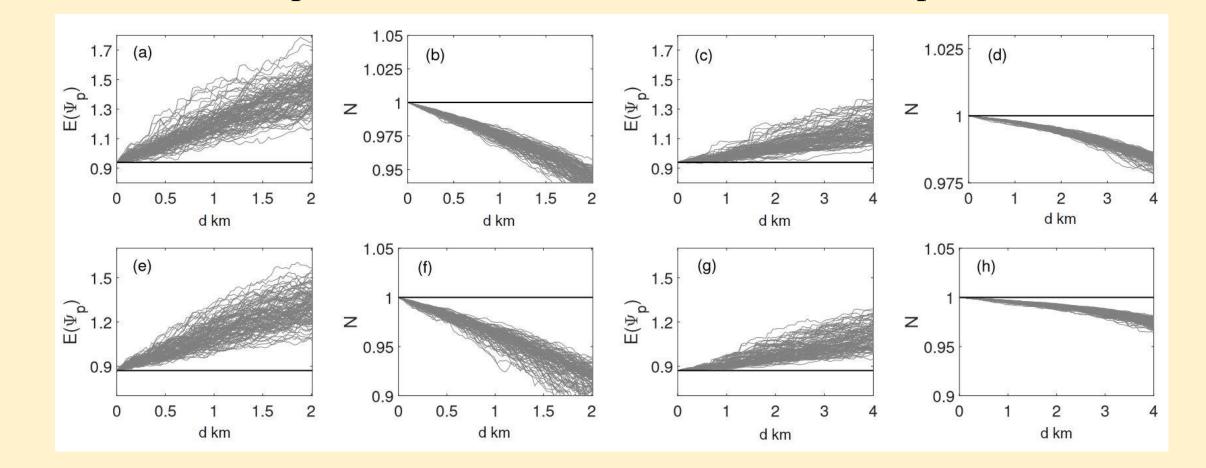
$$\Psi(r,\theta) = \sum_{j,m} c_{jm} \Psi_{jm}(r,\theta) = \sum_{j,m} d_{jm} \psi_j(r;q) \Theta_{\rm m}(\theta)$$

We define  $(j,m)^{\text{th}}$  matrix entry of D as  $(D)_{j,m} = d_{jm}$  and  $\Lambda = D^{\dagger}D$ . The radialangular entanglement  $E(\Psi)$  of  $\Psi(r, \theta)$  is defined through the eigenvalues of  $\Lambda$  as  $E(\Psi) = -\sum_i \lambda_i \log \lambda_i.$ 

### **Observations**

- For any mode superposition  $\Psi(r, \theta)$ , the radial-angular entanglement is invariant under transverse plane rotation.
- For any mode superposition  $\Psi(r, \theta)$  with constituent modes of distinct j and m indices, the radial-angular entanglement in  $\Psi(r, \theta)$  is invariant on passage through any symmetric first-order optical system.
- For any mode superposition  $\Psi(r, \theta)$ , with constituent modes having a fixed j index and varying m index, the radial-angular entanglement in  $\Psi(r, \theta)$  is invariant on passage through a symmetric first-order optical system.
- Free propagation can in general create, preserve, or destroy radial-angular

100 samples for n = 1,3, 5, and 7 and  $C_n^2 \approx 10^{-12} \text{ m}^{-2/3}$  and  $10^{-13} \text{ m}^{-2/3}$ , respectively, for  $\Psi^{1n}$  given in Eq. (1), for d up to 1 km and 2 km, respectively; (c) same, but for n = 1, 2, 3, and 4 and  $C_n^2 \approx 10^{-14} \,\mathrm{m}^{-2/3}$ , and for d up to 5 km for  $\Psi^{1n}(r,\theta)$ . The dark line in each of these frames shows S for the respective  $\Psi^{1n}(r,\theta)$  in the absence of atmospheric turbulence.

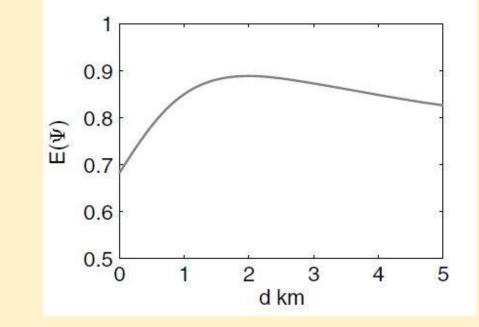


**Fig. 3.** (a), (c)  $E(\Psi_p)$  against distance of propagation d for  $\Psi^{1n}(r, \theta)$  with n = 2 [Eq. (1)] and  $C_n^2 \approx 10^{-13} \text{ m}^{-2/3}$  and  $10^{-14} \text{ m}^{-2/3}$ , respectively, for 100 samples, for d up to 2 km and 4 km; (b), (d) N against distance of propagation d for the samples in (a) and (c), and corresponding strengths of turbulence; (e), (g) [and (f), (h)] same, but for  $\Psi^{1n}(r, \theta)$  with n = 3 for the respective strengths of turbulence. In all the frames, the dark line shows either  $E(\Psi)$  or N(=1)in the absence of atmospheric turbulence.

# Conclusion

To conclude, we have studied radial-angular entanglement in LG mode superpositions with respect to symmetric first-order optical systems. The role of the Gouy phase in this regard has been brought out. We have seen examples of LG mode superpositions for which a symmetric first-order optical system preserves the radial-angular entanglement. The examples studied suggest that atmospheric turbulence can alter the radial-angular entanglement in a mode superposition on propagation through atmospheric turbulence. We have illustrated through examples how the defined witness of radial-angular entanglement can be effective as a free space signaler. This suggests that classical optic entanglement can indeed be of practical value, particularly in the context of free space optical communication.

#### entanglement in a mode superposition $\Psi(r, \theta)$ .



**Fig. 1.**  $E(\Psi)$  against d for the threemode superposition  $\Psi(r, \theta) =$  $\frac{1}{\sqrt{3}}\Psi_{00}(r,\theta) + \frac{1}{\sqrt{3}}\Psi_{11}(r,\theta) + \frac{-i}{\sqrt{3}}\Psi_{21}(r,\theta) .$ 

#### References

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